A HEAT TRANSFER TEXTBOOK

FIFTH EDITION

SOLUTIONS MANUAL FOR CHAPTER 1

by JOHN H. LIENHARD IV *and* JOHN H. LIENHARD V

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- 1.1 Given the wall shown:
	- Find: The temperature distribution in the wall, and suggest any assum-
ptrons that might be
made to simplify analyous of the wall

With reference to Example 1.2 we write:

GOP

102

 $k = 237$ W/m

 $35 W$

corkboard

 $k = 0.04$

 $\frac{1}{2}$ $\frac{1}{2}$ $\frac{1}{2}$ 0.06 m

 $k = 0.12 \frac{W}{mL}$

 $0.05m$

1.2 Verify that Newton's Law of cooling, $-\frac{dT_{body}}{dt} \sim (T_{body} - T_{\omega})$, is equivalent to $\Omega \sim (\text{T}_{\text{body}} - \text{T}_{\infty})$.

We know that:
$$
Q = \frac{dU_{body}}{dt} = \frac{d[oc(volume of body) (T_{body} - T_{ref})]}{dt}
$$

\n
$$
= ocV \frac{dT_{body}}{dt}
$$
\n
$$
= \frac{dT_{body}}{dt} = \frac{Q}{\rho cV} \sim \frac{(T_{body} - T_{av})}{\rho cV}
$$

Thus the relations are equivalent if pcV is constant.

5000 W/m² are transfered through a 1 cm slab whose cold side is held at -40°C. Find ΔT for each of the 7 materials tabled 1.3 below. Discuss the results.

 $q = k \frac{\Delta T}{L}$, where I is the thickness of the slab.

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1.4 Explain why the heat diffusion equation, $\frac{\partial^2 T}{\partial z^2} = \frac{1}{\alpha} \frac{\partial T}{\partial t}$, shows

that in transient conduction T depends on $\alpha = k/\rho c$, but solutions of steady conduction problems involve only k.

In a transient problem (one in which the local temperature of a body changes) some of the heat flow is either stored in, of drawn from, the body. Consequently, both k (which determines the through-flow of heat in accordance with Fourier's Law) and the volumetric heat capacity, pc, (which determines the storage), appear in the heat diffusion equation. They happen to appear in the ratio, α . During steady conduction, the local temperature never changes so pc is not involved in the problem.

The rate of change of entropy of the universe is positive as a result of the process, as the Second Law of Thermodynamics requires it to be.

This problem is exactly the same as 1.5 except the heat transfer rate is higher:

$$
Q = 0.14(.093) \frac{70}{0.1} = 9.114 W
$$

SÔ.

$$
\dot{s}_{res. \#1} = \frac{-0.03038}{-0.03038}, \ \dot{s}_{res, \#2} = \frac{+0.03963}{-0.03963}, \ \dot{s}_{\text{rod}} = \underline{0}
$$

Replace the thermal reservoirs in 1.6 with adiabatic walls. 1.7 If $pc = 1133.4 \text{ kJ/m}^2C$,

a) the final equilibrium temperature, T_f , of the slab; Find: b) AS for the process; c) Is the Second Law satisfied? a.) The final temperature is obviously just a simple mean:

$$
T_f = \frac{27 + (-43)}{2} = -\frac{8 \cdot C}{2}
$$

b.) For a solid, the specific entropy is S-S_{ref} = cln T /T_{re}f Thus we may write at any section of the slab:

 $S_{final} - S_{initial} = c ln T_f/T_i$ $=-c \ln T_i/T_f$ Thus: $\Delta S = \rho A \int_{x=0}^{x=0.1m} (S_f - S_i) dx = -\rho C A \int_{0}^{0.1} (\ln T_i/T_f) dx$

However the initial linear temperature profile is $T_3 = 300^{\circ}K$ - $(700 °K/m)x$; thus $dx = -dT_1/700$ and

- $\Delta S = \frac{\rho C \hat{A}}{700} \int_{300}^{230} (i \pi \tau_i / \tau_f) d\tau_i = 0.1506 [230 \text{ in } \frac{230}{273 8} 230 300 \text{ in } \frac{300}{265} + 300]$
= 0.03081 $\frac{kJ}{\delta K}$
	- c.) The net change entropy in this spontaneous, irreversible process is positive as the second Law requires.

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1.9 Determine the total heat transfer in Problem 1.8. Plot the net entropy generated as a function of temperature.

The minus sign means that the direction of Q is opposite the x-axis.

The water heater shown, has an external surface 1.11 area of $1.3m^2$. Select an insulating material and specify its thickness, to keep the water from cooling at more than 3°C/hr. Justify neglecting the thermal resistance of the steel casing and the convective layer.

First determine the heat flux: $q = \frac{Q}{A} = \frac{(\text{mass})c}{A} \frac{dT}{dt} = \frac{(100)(4190)3}{1.3(4600)}$ $= 268.6$ W/m²

Then
$$
q = k \frac{\Delta T}{t}
$$
 or $\frac{k}{t} = \frac{268.6}{75-20} = 4.883 \frac{W/m2C}{m}$

Now we look at various insulating materials, and see how large t must be

> for 85% magnesia: $k = 0.068$, $t_{min} = .014$ m for 980 kg/m³ asbestos: $k = 0.14$, $t_{min} =$ for glass wool: $k = 0.04$, $t_{min} = .0082$ m

Let's specify 1-1/2 cm of glass wool. That will be plenty safe-With $q = only 269 W/m²$, the temperature drop through a typical \overline{h} = 10/m²-°C, ΔT = g/h = 27°C. It appears that neglecting h is not justified. We don't really need so much insulation if the air is still around the heater.

And with $q = 269$ W/m², temperature drop through, say, 5 mm of steel is $\Delta T = (q/k) (.005 m) = \frac{269}{54} (.005) = 0.025°C$, which is entirely negligible.

1.12 The two walls shown offer negligible
\nthermal resistance. Find the temperature
\non the L.H.S.
\n
$$
q = h_{LHS}(373 - T_{LHS}) = \sigma (T_{LHS}^4 - T_{RHS}^4)
$$

\n $= h_{RHS}(T_{RHS}^{-293})$
\n $T_{\infty}^{-393 \text{°K}}$
\n T_{LHS}^T
\n T_{RHS}^T
\n T_{RHS}^T

To simplify computation, define $t = T/100$. Then

50 (100)
$$
(3.73-t_L) = 5.67(t_L^4-t_R^4) = 20(100) (t_R^2-2.93)
$$

or

$$
3.73-t_{L} = \frac{5.67}{5000} (t_{L}^{4} - t_{R}^{4}) = 0.4t_{R} - 1.172
$$

Then

$$
t_L = 4.902 - 0.4t_R
$$
 and $t_R = 12.255 - 2.5t_L$

and

$$
t_{\rm L} = 3.73-0.001134 \left(t_{\rm L}^{4} - [12.255-2.5t_{\rm L}]^{4}\right)
$$

which must be solved by trial and error

 $T_R = 42.5^{\circ}C$

and

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Develop conversion factors for α , g, ρ , σ , F_{1-2} , S , c: α : $1 \frac{m^2}{s} = 1 \frac{m^2}{s} \left(\frac{ft}{3048 m} \right)^2 = 10.764 \frac{ft^2}{s}$ q: $1 \frac{W}{m^2} = 1 \frac{J}{m^2s} \left(\frac{.3048 \text{ m}}{\text{ft}} \right)^2 \left(\frac{0.00094783 \text{ Btu}}{J} \right) \frac{3600 \text{ s}}{\text{hr}} = 0.317 \frac{\text{Btu}}{\text{ft}^2 \text{hr}}$ p: $1 \frac{\text{kg}}{\text{m}} = 1 \frac{\text{kg}}{\text{m}^3} (\frac{2.2046 \text{ lbm}}{\text{kg}}) (\frac{.3048 \text{ m}}{\text{ft}})^3 = 0.06243 \frac{\text{lbm}}{\text{ft}^3}$ $\sigma: \frac{1}{m^2 \cdot \kappa^4} = 1 \frac{W}{m^2 \cdot \kappa} \cdot \frac{317 B t u/f t^2 h r}{W/m^2} \left(\frac{K}{1.8 \cdot R}\right)^4 = 0.03020 \frac{B t u}{f t^2 h r^2 R^4}$ $F_{1-2}: 1 = 1$ (since F_{1-2} is dimensionless, the conversion factor is unity.) - \hat{s} : $\frac{J}{kg \text{ mole}^2 K}$ = 1 $\frac{J}{kg \text{ mole}^2 K}$ ($\frac{.0094783Btu}{J}$) ($\frac{kg \text{ mole}}{2.204621bm \{621}$) ($\frac{1}{2.204621bm \{621}$) $X(\frac{8K}{1.89R})$ $= 0.00023885 \frac{Btu}{lb_m^{\text{mole}}R}$ c: $1 \frac{kJ}{kg^9C} = 1 \frac{kJ}{kg^9C} (\frac{0.94783Btu}{kJ})(\frac{kg}{2.204621b_m})(\frac{°C}{1.8°F}) = 0.23885 \frac{Btu}{1b_n^{~9}F}$

1.14 Given 3 infinite parallel black
\n
$$
\frac{100^{\circ}C}{2}
$$
\n273^oL
\n
$$
\frac{100^{\circ}C}{2}
$$
\n
$$
\frac{100^{\circ}C}{2}
$$
\nby radiatron and the plates
\nare open 4 are open 4 .\n
\n
$$
q = \sigma(\tau_1^4 - \tau_2^4) = \sigma(\tau_2^4 - \tau_3^4)
$$
 so $2\tau_2^4 = \tau_1^4 + \tau_3^4$
\n
$$
\tau_2 = \sqrt{0.5(373)^4 + (273)^4} = 334.07^{\circ}K = \frac{61.07^{\circ}C}{4}
$$

Same as 1.15 but now there are four plates and we want two 1.15

$$
4 = \underbrace{\underbrace{\sigma(\tau_1^4 - \tau_2^4)}_{T_2} = \underbrace{\sigma(\tau_2^4 - \tau_3^4)}_{T_3} = \underbrace{\sigma(\tau_3^4 - \tau_4^4)}_{2\sigma_3} \cdot \underbrace{\underbrace{\tau_4^4 - \tau_5^4}_{T_4} = \underbrace{\tau_5^4 - \tau_6^4}_{T_5} \cdot \underbrace{\tau_6^4 - \tau_7^4}_{T_6} = \underbrace{\tau_7^4 - \tau_8^4}_{T_7} = \underbrace{\tau_8^4 - \tau_9^4}_{T_8} = \underbrace{\tau_9^4 - \tau_9^4}_{T_9} = \underbrace{\tau_9^4 - \tau
$$

 $T_t^4 - T_t^4 = T_t^3 - T_a^4$ Than $T_2 = \sqrt[4]{T_1^4 + T_4^4 - T_3^4} = 348.53^{\circ}K = 15.53^{\circ}C$ T_{2}

1.16 Consider the conduction-radiation configuration shown:

Write the relation:

$$
\frac{q_{rad}}{q_{cond}} = f(N, \text{e}) \equiv \frac{T_h}{T_c}
$$

 W_2 ($k \approx 0.1$ $L = 10cm$ $\overline{1}$ = 100°

a dimensionless group to be determined

and plot it for q_{rad}/q_{cond} , = 0.8, 1, and 1.2. Identify
the given operating point and find T_h .

$$
\frac{q_{\text{rad}}}{1-\frac{1}{2}} = \frac{q(T_h - T_c)}{r_h/L} = \frac{q(T_h/T_c)}{r_h/L} = \frac{q(T_c)}{r_h/L} = \frac{1}{N} \frac{Q^2 - 1}{Q - 1}.
$$

If $q_{rad}/q_{cmd} = 1$, then $N = (\bigotimes^{d} -1) / (\bigotimes -1)$ which is plotted below:

12

Problem 1.18 A small instrument package is released from a space vehicle. We can approximate it as a solid aluminum sphere, 4 cm in diameter. The sphere is initially at 303 K and it contains a pressurized hydrogen component that will condense and malfunction at 30 K. If we approximate outer space to be at 0 K, how long will the instrumentation package function properly? Is it legitimate to use the lumped-capacity method? **Solution**

$$
\int c \sqrt{\frac{dT}{dt}} = -\frac{\pi A T^4}{2}
$$

\n
$$
\int_{303}^{30} \frac{dT}{T^4} = \int_{0}^{t} \frac{\pi A}{2} dt \qquad \sigma \qquad \frac{1}{3T^3} \int_{303}^{30} = \frac{3\pi}{2} t
$$

\nThus:
\n
$$
t = \int_{307}^{0} \frac{cR}{2} \left[\frac{1}{30^3} - \frac{1}{303^3} \right] = \frac{2707 (965)(0.02)}{9(5.61)10^{-25}} [0.000037] = 3.55 \times 10^6 s
$$

\n= 331hr
\n
$$
\int_{0}^{20} \frac{dT_1^3}{2} = \frac{5.67(3.03)10^{-8}}{9(5.61)10^{-25}} [0.000037] = 3.55 \times 10^6 s
$$

\n= 331 hr
\n
$$
= \frac{41 \text{ days}}{9.0000126} = 0.0000126 \qquad \text{4} \qquad \text{(no problem)}
$$

function.
\n
$$
\frac{\partial^{2}T}{\partial x^{3}} = \frac{1}{\alpha} \frac{\partial T}{\partial t} \Rightarrow \frac{d^{2}T}{dx^{2}} = 0
$$
 In this case. Incomplete, twice and get $\frac{T = C_{1} + C_{2} \times \sqrt{1 - T_{1}}}{T = C_{1} + C_{2} \times \sqrt{1 - T_{1}}$ so $C_{1} = T_{1}$ and at $x = 1$, $T = T_{r}$ so $C_{2} = (T_{r}$ is T_{1}),
\n
$$
\frac{T = T_{r}}{1 + T_{1}}
$$

1.21 Consider the wall shown.

This means that q may be evaluated precisely in plane steady heat transfer if we use the mean value of k.

122 Find k for the wall shown:
\n
$$
\frac{122}{1000} = \frac{1}{4} \text{Hrough} \cdot \text{ln} \cdot \text{ln
$$

$$
q = \frac{2(100 - 99.3)}{0.02} = \frac{1070 \frac{W}{m^{2}}}{T_{j}} = \frac{5(25 - T_{j})}{0.04}
$$

T_j = -0.008(1070) + 25 = 16.44°C

$$
q = 1070 = \frac{4(16.44 - T_r)}{0.04} = 1644 - 100 T_r
$$
, $T_r = 5.74\% \rightarrow$

An aluminum beverage can (12 cm high and 6 cm in diameter) is 1.24 initially at refrigerater temperature--say 3° C. It is placed
in the kitchen at--say--25°C. If $F = 13.5$ W/m²-°C, how long
will it take for the beverage to reach 15°C? (This would be a 12 oz , $can.$)

Additional assumptions: The beverage has the properties of water. The aluminum offers no thermal capcitance. The bottom of the can resting on the table is insulated.

The beverage along the inner walls of the can will warm up first and rise, then stir in with the cooler liquid toward the middle. That action, not conduction, will keep the temperature close to a mean value as it warms. The Biot number is thus meaningless. But the liquid will nevertheless remain close to a constant temperature and the lumped capacitance assumption will be valid.

$$
\mathbf{T} = \frac{\rho c V}{\overline{\mathbf{h}}\mathbf{A}} = \frac{1000 (4190) \pi (0.03)^{2} (0.12)}{13.5 [\pi (.03)^{2} + \pi (.06) (0.12)]} = 4138 \text{ sec.}
$$

We want $\frac{T-T_{\infty}}{T_i-T_{\infty}} = \frac{15-25}{3-25} = 0.4546$, so eqn. (1.21) gives: $0.4546 = e^{-t/4138}$

Thus the cooling time, t , = 3263 sec or 54.4 min.

1.25 Find the far temperature of the 0.1 m wall, desinated as T_a . Assume that both walls are black

$$
-k \frac{\Delta T}{L} = \sigma \left(\left[127 + 273 \right]^4 - \left[27 + 273 \right]^4 \right)
$$

$$
17.5 \frac{T_a^{-127}}{0.1} = 5.67 \cdot (4^4 - 3^4) = 992.25
$$

A 1 cm diam, 17 carbon steel, sphere, initially at 2002, 1.26 is cooled by natural convection with air at 20%. In this case h is not independent of temperature. Instead, \overline{h} = 3.51(ΔT °C)^{1/4} W/m². °C. Plat Tsphere as a function of t. Verify the lumped capacity assumption.

$$
\frac{\partial(\Gamma - T_{\infty})}{\partial t} = -\frac{\bar{h} \Delta}{\rho c V} (\tau - T_{\infty}) = -\frac{3.51}{\rho c (R/3)} (\tau - T_{\infty})^{5/4}
$$

 \circ \circ

50

$$
\int_{T=2\infty}^{T} \frac{d(T-T_{\infty})}{(T-T_{\infty})^{3/4}} = -\int_{0}^{1} \frac{10.53 dt}{\rho cR} = -\frac{10.53}{\rho cR} +
$$

$$
-\frac{1}{4}\left[\frac{1}{(T-20)^{1/4}}-\frac{1}{180^{1/4}}\right]=- \frac{10.53}{7801(473)(0.005)}t=-0.000571t
$$

$$
t = 438 \left[\frac{1}{(T-20)^{1/4}} - 0.273 \right] \leftarrow
$$

1.27 A 3cm diam., black spherical header is kept at 1100°C.
\nIt radiates through an evacuated annulus, to a surr-
\nounding spherical shell of Nichrome
$$
\overline{L}
$$
. The shell has a
\n3 cm LD and is 0.3 cm thick. It is black on the in-
\nside and it is held at 25°C on the outside: Find:
\na) The temperature of the inner wall of the shell and b) the
\nheat transfer, Q . (Treat the shell as a plane wall.)

$$
T_h = 1373^{\circ}K
$$

\n $T_i = ?$
\n $Q_{rad} = Q_{cond}$
\n $T_i = ?$
\n $Q_{rad} = Q_{cond}$
\n Q_{rad}
\n Q_{rad}
\n Q_{rad}
\n Q_{rad}
\n Q_{rad}
\n Q_{cond}
\n Q_{rad}
\n Q_{cond}
\n

Th

 \bar{t}

Then:
\n
$$
Q = Q_{rad} = 5.67 \times 10^{-8} [\pi (0.03)^{2}] (1373^{4} - 304.85^{4})
$$

\n
$$
= 568 \text{ W}
$$
\n
$$
= 568 \text{ W}
$$
\n
$$
= 568 \text{ W}
$$
\n
$$
= 568 \text{ W}
$$

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Problem 1.28 The sun radiates 650 W/m² on the surface of a particular lake. At what rate, \dot{m} mm/hr, would the lake evaporate away if all this energy went to evaporating water? Discuss as many other ways this energy can be distributed as you can think of. (h_{fg} for water is 2,257,000 J/kg.) Do you suppose much of the incident radiation goes to evaporation?

> Q_{rad} = Q_{latent} 650 J/m²-s = ρ_f kg/m³ [\dot{m} /1000(3600) m/s] h_{fg} J/kg $=$ [997(2,257,000)/3,600,000] m

So the maximum possible \dot{m} would be 1.04 mm/hr

Other places Q_{rad} could go, include:

- (a) Much of the visible portion of Q_{rad} reflects off the surface. (The visible part of solar radiation is quite large so this could be important. See Fig. $11.2.)$
- Some of the visible and ultra-violet parts of Q_{rad} will be transmitted (b) and/or absorbed below the surface.
- (c) The infrared portion, which is absorbed at the surface, can go three ways:
	- Natural convection will return some to the air.
	- Some will conduct to the colder water below.
	- Some will reradiate to the sky at night.

Thus, the surface cannot get very hot, and the rate of evaporation should be far less than our calculated value of 1.04 mm/hr.

It is proposed to make picnic cups, 0.005 m thick, of a new plastic 1.29 for which k=k_o(1+aT²) where T is expressed in °C, k_o=0.15 W/m-°C, and $a=10^{-4}$ °C⁻². We are concerned with thermal behavior in the extreme case in which T=100°C in the cup and 0°C outside. Plot T against position in the cup wall and find the heat loss, q.

In this case:
$$
q = k_o(1+aT^2)\frac{dT}{dx} = constant
$$

$$
\frac{q}{k_o} \int_{0}^{L=0.005} dx = \int_{T(k=0)}^{T(k=L)} dT
$$
 where $T(x=L)=100$
 $T(x=0)=0$

or

$$
\frac{4L}{k_o} = \left[T + \frac{a}{3} T^3 \right]_o^{100}
$$

$$
30: \qquad q = \frac{0.15}{0.005} \left[100 + \frac{10^{-4}}{3} 10^{6} \right] = \frac{4000 \text{ W/m}^2}{}
$$

Had we integrated only to x denotes
$$
-
$$
 not to L, then:
\n $q = 4000 = \frac{k_0}{x} [T(x) + \frac{q}{3} T(x)]$

OY

 $X = \frac{0.15}{4000} \left[T(x) + \frac{10^{-4}}{9} T^3(x) \right]$

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A disc-shaped wafer of diamond IIb is the target of a very high intensity laser. The disc is 5 mm in diam. and 1 mm deep. The flat side is pulsed intermittently with 10^{10} W/m² of energy for one microsecond. It is then cooled by natural convection from that same side until the next pulse. If \overline{h} =10 W/m²-°C and T_{∞} =30°C, plot T_{disc} as a function of time for pulses that are 50 sec apart and 100 sec apart. (Note that you must determine the temperature the disc reaches before it is pulsed, each time.)

First evaluate
$$
\frac{1}{\pi} = \frac{\overline{h}A}{\rho c V} = \frac{10 \frac{\pi}{4} (0.005)^2}{3250(510) \frac{\pi}{4} (0.005)^2 (0.001)} = \frac{1.963 (10)^4}{0.03254} = \frac{1}{165.8}
$$

The disc receives 10¹⁰
$$
\frac{J}{m^2-s}
$$
 $\frac{J}{4}$ (0.005) $m^2 \times 10^{-6}$ s = 0.1964 $\frac{J}{puls}$
so 1¹ s $\frac{J}{2}$ $\frac{m}{m} = 6.034 \frac{c}{muls}$
 $T_i - T_f = \frac{0.1964 J}{\rho cV J/\tau} = 6.034 \frac{c}{puls}$

The cooling will be as described by eqn. (1.21)
\n
$$
\frac{T-T_{\alpha}}{T_{\hat{L}}-T_{\alpha}} = e^{-\frac{t}{165.8}} = e^{-\frac{t}{165.8}}
$$
\n
$$
s_{\omega} = \frac{T_{\hat{L}}-30}{T_{\hat{L}}-(30-6.034)} = e^{-\frac{50 \text{ (165.8)}}{165.8}}
$$

We solve this for T_f and ge + $T_f = 47.17$ for SO securiterial
Then using eqn. (1.21) we plot: $T_f = 37.29$ for 100 securiterial T_{disc} ℃ 1 pulse/50 s 50 47.17

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- 1.31 An old incandescent 60W light bulb is roughly a 0.06m diam. sphere. Its surface temp. is 115°C. The average heat transfer coefficient outside the bulb is 8.2W/mºK
	- (a) Show that the peak radiation from the glass to the room has a near-infrared wavelength.
	- (b) What is the heat loss from the glass if $\varepsilon_{glass} = 0.94$?
	- (c) How much heat transfer remains to occur by direct radiation from the filament through the glass? Most of that energy is not in the visible spectrum. These bulbs were very inefficient.)
- a.) Wien's Law, eqn. 1.29: μ at max e_b = 2898(115 + 273) $= 7.47$ um

This places the radiation: squarely in the near-infrared $(See Table 1.2)$

b.) Q from glass =
$$
\overline{h}A\Delta T + \sigma A(T^4_{\text{bulb}} - T^4_{\text{CD}})
$$
 ϵ_{glass}
= $7\pi (0.06)^2 (115-25) + 5.67 (10)^8 \pi (388^4 - 290^4) 0.34$
= 7.125 W + 9.397 W = 16.52 W

c.) Percent of direct radiation from filament $=$ (115 - 16.52)100/115 = 85.6 percent

1.32 How much entropy does the light bulb in Problem 1.31 produce?

$$
\frac{5}{5} = \frac{5}{5}
$$
energy + bulb + 5
\n0 since energy of at steady L = $\frac{a_{\text{to room}}}{T_{\text{room}}}$
\ne horizontal work
\n= 0 + 0 + $\frac{115 \text{ J/s}}{(273 + 25) \text{ K}} = 0.4637 \frac{\text{J}}{\text{Ksec}}$

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 2.2

Air at 20°C flows over one side of a thin metal sheet $(\overline{h} = 10.6 W/m^2 - C)$. $1.33 -$ Methanol at 87°C flows over the other side $(\overline{h} = 141 \text{ W/m}^2 - \text{°C})$. The metal functions as an electrical resistance heater, releasing 1000 w/m^2 . Calculate: a) the heater temperature, b) the heat transfer from the methanol to the heater, and c) the heat transfer from the heater to the air.

$$
\frac{\overline{h} = 10.6}{\overline{h} = 141}
$$
\n
$$
\frac{(\overline{r}_{a} = 202)}{(\overline{r}_{m} = 812)}
$$
\n
$$
= 141(81 - \overline{r}_{plane}) + 1000 = 10.6(\overline{r}_{plate} - 20)
$$
\n
$$
\frac{\overline{r}_{plate}}{11}
$$
\n
$$
\frac{\overline{r}_{plate}}{11}
$$
\n
$$
= 88.91°C
$$

Then 9_{meth to heat} = 141(87-88.5) =
$$
= 269.5 \text{ W/m}^2
$$

4_{heat} to an = 10.6(88.91-20) =
$$
\frac{730.5 \text{ W/m}^2}{}
$$

A black heater is simultaneously cooled by 20°C air (with \overline{h} =14.6 W/m²-°C) 1.34 and by radiation to a parallel black wall at 80°C. What is the temperature of the first wall if it delivers 9000 w/m^2 .

$$
3000 = 14.6(T_{w} - [20 + 213]) + 5.67(10)^{8} (T_{w}^{4} - [80 + 213])^{4})
$$

 $O(r)$

$$
T_w^4 = 24.91(10)^{10} - 0.02575(10)^{10}T_w
$$

Solve by byn 12,36(10)¹⁰ 8.52(10)¹⁰
\n
$$
600^\circ K
$$
 12,36(10)¹⁰ 9.52(10)¹⁰
\n550 9.15 10.81
\n560 9.835 10.55
\n565 10.19 10.72
\n568 10.41 10.34
\n567 10.34 10.37
\n $60^\circ K$ 2294°C
\n $60^\circ K$ 2294°C

Problem 1.35: A 250 mL (8.3 oz.) aluminum beverage can is taken from a 3°C refrigerator and placed in a low humidity, 25°C room (\hbar = 7.3 W/m²K). The 53.3 mm diameter by 112 mm high can is placed on an insulated surface. How long will it take to reach 12°C? Assume that emittance of the can is very low, so thermal radiation is negligible. Discuss your other approximations.

Solution: The beverage has about the same properties as water ($\rho = 999.9 \text{ kg/m}^3$, $c =$ 4200 J/kg K). We can neglect the heat capacitance of the thin aluminum can. We can also assume that the liquid in the can will circulate under buoyancy, so internal temperature gradients will remain small. Then a lumped capacitance solution applies.

The surface area for heat transfer is the top and sides of the can:

 $A = \pi D^2/4 + \pi D L = \pi [(53.3)^2/4 + (53.3)(112)] \times 10^{-6}$ m² = 0.02099 m²

and the mass of liquid is:

$$
m = \rho V = (999.9 \text{ kg/m}^3)(250 \times 10^{-6} \text{ m}^3) = 0.2500 \text{ kg}
$$

The time constant is

$$
\mathbf{T} = \frac{mc}{\bar{h}A} = \frac{(0.2500)(4200)}{(7.3)(0.02099)}
$$

$$
=6953
$$

Using eqn. (1.22) with $T_1 = 3^\circ \text{C}$ and $T_\infty = 25^\circ \text{C}$, we solve for the time at which $T = 12^\circ \text{C}$.

$$
\frac{12 - 25}{3 - 25} = e^{-t/6853}
$$

$$
0.5909 = e^{-t/6853}
$$

$$
t = -\ln(0.5909) \cdot (6853 \text{ s}) = 3605 \text{ s} = 1 \text{ h } 5 \text{ s}
$$

Problem 1.36: A thin sheet is a resistance heater, parallel with 3 cm slabs of cast iron on either side, in an evacuated cavity. The heater, which releases 8000 W/m², and the cast iron, are very nearly black. The outside surfaces of the cast iron slabs are held at 10°C. Find the temperatures of the heater and the inside of the slabs.

A black wall at 1200°C radiates to the left side of a parallel slab 1.37 of 316 stainless steel, 5 mm thick. The right side of the slab is to be cooled convectively and is not to exceed 0°C. Suggest a convective process that will achieve this.

So: 2.215 $\frac{0+10}{0.02}$ = \overline{h} (15-0), \overline{h} incipient melting = 73.8W/m². ez
if Touter < 0°C, \overline{h} will be lower. There fere \overline{h} must to raised
to make melting progress. to make melting progress.

Problem 1.39 At what minimum temperature does a black radiator have its maximum monochromatic emissive power in the visible wavelength range? Look at Fig. 10.2; then describe the difference between what you might see looking at this object in comparison to looking at the sun.

Solution In accordance with eqn. (1.28), and using $\lambda_{\text{max.visible}} = 0.00008 \text{ cm}$,

0.00008 T = 0.2898 cm K. Therefore $T = 3623$ K

The sun radiates at about 5777 K (see Fig. 10.2). This is a substantially higher temperature. It also delivers its maximum e_{λ} at a much smaller wavelength – one at the lower end of the visible range.

Problem 1.40 The local heat transfer coefficient during the laminar flow of fluid over a flat plate of length L is equal to $F/x^{1/2}$, where F is some function of fluid properties and the flow velocity. How does the average heat transfer coefficient compare with $h(x=L)$ if x is the distance from the leading edge of the plate?

Solution: We use the definition of the average to get:

$$
\bar{h} = \frac{1}{L} \int_0^L h \, dx = 2 \frac{F}{L} x^{1/2} \Big|_0^L = 2 \frac{F}{\sqrt{L}} = 2 h(x = L)
$$

Therefore, the average heat transfer coefficient = $\overline{h} = 2 h(x = L)$

Problem 1.41 An object is initially at a temperature higher than its surroundings. We have seen that many kinds of convective processes will bring the object into equilibrium with its surroundings. Describe the characteristics of a process that will do so with the smallest net increase in the entropy of the universe.

Solution Entropy is *not* a path function. *Any* process connecting the initial to the final states will yield the same increase of entropy.

26

4.42 A 250°C cylindrical copper billet, 4 cm in diameter and 8 cm long, is cooled in air at 25°C. The heat transfer coefficient is 5 W/m^2 -°C. Can this be treated as lumped capacitance cooling? What is the temperature of the billet after 10 minutes?

Volume =
$$
\frac{\pi}{4} (0.04)^2 (0.08) = \frac{0.0001005 \text{ m}^3}{9.0001005 \text{ m}^3}
$$

Area = 0.01257 m²

Base bi on V/A = 0.008 m (We could also have used R/Z = 0.010
or some other possibilities on the
same order of magnitude.)
Then Bi =
$$
\frac{\overline{h}V/A}{k} = \frac{S(0.008)}{391} = 0.000102 \ll \frac{1}{2}
$$
. It is OK to use
lumped capacity.

 $\mathbf{T} = \frac{\rho cV}{hA} = 384(8954) \frac{0.008}{5} = 5501 sec$
So: $\frac{T(\text{I0} \text{ min}) - 25}{250-25} = e^{-600 sec/550sec}$ $T(\text{I0} \text{ min}) = 226.8^{\circ}$

PROBLEM 1.43: The diameter of the sun is roughly 1,391,000 km, and it emits energy as if it were a black body at about 5772 K. Determine the rate at which it emits energy. Compare this with the known value. How much energy does the sun emit per year? $[1.21 \times 10^{34}$ J/y]

Older versions of AHTT used 5777 K for the solar temperature and 1,387,000 km for the solar diameter. The results are nearly identical.

SOLUTION

The radiative power *emitted by* the sun is

$$
Q_{\text{sun}} = \pi D_{\text{sun}}^2 \sigma T_{\text{sun}}^4
$$

= $\pi (1.391 \times 10^9)^2 (5.670374 \times 10^{-8}) (5772)^4$
= 3.826×10^{26} W

With the SI prefixes in Table B.1, we could instead write $Q_{sun} = 382.6$ YW ("yottawatts"). This value is a nearly exact match to the 2015 standard value of 3.828×10^{26} W [See comment below].

The annual energy is

$$
E_{\text{sun}} = (365.25)(24)(3600)(3.826 \times 10^{26}) = 1.207 \times 10^{34} \text{ J/y} = 1.207 \times 10^{3} \text{ QJ/y}
$$

Comment. The International Astronomical Union provides the solar data used here: Prša et al., "Nominal Values for Selected Solar and Planetary Quantities: IAU 2015 Resolution B3," Astronomical Journal 152:41, 2016, doi:10.3847/0004-6256/152/2/41.

Problem 1.44: Room temperature objects at 300 K and the sun at 5772 K each radiate thermal energy; but Planck's law, eqn. (1.30), shows that the wavelengths of importance are quite different.

- a) Find λ_{max} in micrometers for each of these temperatures from Wien's Law, eqn. (1.29).
- b) Using a spreadsheet or other software, plot eqn. (1.30) for $T = 300$ K as a function of wavelength from 0 to 50 µm and for $T = 5772$ K for wavelengths from 0 to 5 µm.
- c) By numerical integration, find the total area under each of these curves and compare the value to the Stefan-Boltzmann law, eqn. (1.28). Explain any differences.
- d) Show that about 1/4 of the area under each curve is to the left of λ_{max} (in other words, 3/4 of the energy radiated is on wavelengths greater than λ_{max}).
- e) What fraction of the energy radiated by the 300 K surface is carried on wavelengths less than 4 µm? What fraction of the energy radiated by the 5772 K surface is on wavelengths greater than 4 µm?

Earlier versions of AHTT used 5777 K for the solar temperature. The results are nearly identical.

SOLUTION.

a)

$$
\lambda_{\text{max}} = \frac{2987.77 \text{ }\mu\text{m} \cdot \text{K}}{T \text{ K}} = \begin{cases} 9.9592 \text{ }\mu\text{m} & \text{at } 300 \text{ K} \\ 0.5176 \text{ }\mu\text{m} & \text{at } 5772 \text{ K} \end{cases}
$$

b) The plotting and integration can be done in various ways depending upon what software is used. The results in Fig. 1 are from an Excel spreadsheet with a step size of 0.2 µm at 300 K and of 0.02 µm at 5772 K.

(a) 300 K

(b) 5772 K

Figure 1. Plots of Planck's law at two temperatures

c) Using the values from Excel, a trapezoidal rule integration gives the area under the curve:

Integrated area =
$$
\begin{cases} 445.2 \text{ W/m}^2 & \text{at } 300 \text{ K} \\ 6.261 \times 10^7 \text{ W/m}^2 & \text{at } 5772 \text{ K} \end{cases}
$$

The Stefan-Boltzmann law yields (with $\sigma = 5.670374 \times 10^{-8}$ W/m² K⁴)

$$
\sigma T^4 = \begin{cases} 459.3 \text{ W/m}^2 & \text{at } 300 \text{ K} \\ 6.294 \times 10^7 \text{ W/m}^2 & \text{at } 5772 \text{ K} \end{cases}
$$

At 5772 K, the integrated value is 99.48% of the Stefan-Boltzmann law, and at 300 K, it is 96.93%. The principal reason that these values are low is that energy is also radiated at wavelengths higher than the range of integration.

d) By integrating up to λ_{max} from part a),

$$
\frac{1}{\sigma T^4} \int_0^{\lambda_{\text{max}}} e_{\lambda,b}(T) \, d\lambda = \begin{cases} 28.4\% & \text{at } 300 \text{ K} \\ 29.6\% & \text{at } 5772 \text{ K} \end{cases}
$$

These values are bit more than 1/4 of the total energy, but are often stated as "about 1/4".

e) Similar integrations show that a 300 K surface radiates only 0.33% on wavelengths below 4 µm and that a 5772 K surface radiates 99.0% on wavelengths less than 4 µm (or 1% on wavelengths above 4 µm). This fact enables the design of materials that selectively absorb or reflect solar energy (see Section 10.6).

Problem 1.45

A crucible of molten metal at 1800°C is placed on the foundry floor. The foundryman covers it with a metal sheet to reduce heat loss to the room. If $\mathcal F$ is 0.4 between the melt and the plate and 0.8 between either the melt or the top of the plate and the room, how much will the heat loss to the room be reduced by the sheet?

SOLUTION. First find the sheet temperature:

$$
q = (0.8) \sigma \left[T_s^4 - (20 + 273)^4 \right] = (0.4) \sigma \left[(1800 + 273)^4 - T_s^4 \right]
$$

This gives $T_{\text{sheet}} = 1575$ °K, so

$$
\frac{q_{\text{with sheet}}}{q_{\text{without sheet}}} = \frac{0.8\sigma(1575^4 - 293^4)}{0.8\sigma(2073^4 - 293^4)} = 0.333
$$

The heat loss is therefore reduced by 66.7% by the shield.

PROBLEM 1.46: Integration of Planck's law, eqn. (1.30) over all wavelengths leads to the Stefan-Boltzmann law, eqn. (1.28). Perform this integration and determine the Stefan-Boltzmann constant in terms of other fundamental physical constants. *Hint:* The integral can be written in terms of Riemann's zeta function, $\zeta(s)$, by using this beautiful relationship between the zeta and gamma functions

$$
\zeta(s)\,\Gamma(s) = \int_0^\infty \frac{t^{s-1}}{e^t - 1} \, dt
$$

for $s > 1$. When s a positive integer, $\Gamma(s) = (s - 1)!$ is just a factorial. Further, several values of $\zeta(s)$ are known in terms of powers of π and can be looked up.

SOLUTION.

(1)
$$
e_b(T) = \int_0^\infty e_{\lambda,b} d\lambda
$$

(2)
$$
= \int_0^\infty \frac{2\pi hc_o^2}{\lambda^5 \left[\exp(hc_o/k_BT\lambda) - 1 \right]} d\lambda
$$

(3)
$$
= \int_0^\infty \frac{2\pi h v^3}{c_o^2 \left[\exp(h v / k_B T) - 1\right]} dv
$$

(4)
$$
= \frac{2\pi k_B^4 T^4}{h^3 c_o^2} \int_0^\infty \frac{x^3}{e^x - 1} dx
$$

We are given

$$
\zeta(s)\,\Gamma(s) = \int_0^\infty \frac{t^{s-1}}{e^t - 1} \, dt
$$

For our case, $s = 4$ and $\Gamma(4) = 3! = 6$. Hence:

(5)
$$
e_b(T) = \frac{2\pi k_B^4 T^4}{h^3 c_o^2} \zeta(4) 3!
$$

(6)
$$
= \frac{12\pi k_B^4}{h^3 c_o^2} \zeta(4) T^4
$$

Zeta is a famous function, and the value at 4 has been established to be:

$$
\zeta(4) = \frac{\pi^4}{90}
$$

Hence:

(7)
\n
$$
e_b(T) = \left(\frac{2\pi^5 k_B^4}{15h^3 c_o^2}\right) T^4
$$
\n
$$
= \sigma T^4
$$

where we have also found the Stefan-Boltzmann constant in terms of fundamental physical constants.